Relativistic electron beams above thunderclouds


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Abstract. Non-luminous relativistic electron beams above thunderclouds have been detected by the radio signals of low frequency ~40–400 kHz which they radiate. The electron beams occur ~2–9 ms after positive cloud-to-ground lightning discharges at heights between ~22–72 km above thunderclouds. Intense positive lightning discharges can also cause sprites which occur either above or prior to the electron beam. One electron beam was detected without any luminous sprite which suggests that electron beams may also occur independently of sprites. Numerical simulations show that beams of electrons partially discharge the lightning electric field above thunderclouds and thereby gain a mean energy of ~7 MeV to transport a total charge of ~10 mC upwards. The impulsive current ~3 × 10⁻² Am⁻² associated with relativistic electron beams above thunderclouds is directed downwards and needs to be considered as a novel element of the global atmospheric electric circuit.

1 Introduction

Sprites (Neubert et al., 2008; Füllekrug et al., 2006; Rakov and Uman, 2003; Sentman et al., 1995; Franz et al., 1990) and gigantic jets (Cummer et al., 2009; Su et al., 2003; Pasko et al., 2002) are luminous streamer (Luque and Ebert, 2009; Pasko et al., 1998) and leader (Riousset et al., 2010; Krebbiel et al., 2008; Pasko and George, 2002) discharges resulting from conventional electrical breakdown in the middle atmosphere. Sprites may also be associated with electron beams resulting from relativistic runaway breakdown within or above thunderclouds (Östgaard et al., 2008; Inan, 2005; Smith et al., 2005; Roussel-Dupré and Gurevich, 1996; Fishman et al., 1994). The newly-recognised runaway breakdown mechanism (Gurevich and Zybin, 2005; Gurevich et al., 1992) inside thunderclouds (Shao et al., 2010; Dwyer et al., 2009) can initiate at ~1/10 of the conventional breakdown electric field threshold (Marshall et al., 1995) and requires a population of energetic seed particles to initiate an electromagnetically radiating non-linear electron avalanche growth (Gurevich et al., 2003). The population of seed particles must be above the runaway threshold of ~10 keV or more, depending on the strength of the electric field, and could originate from ~10⁻¹⁵–10⁻¹⁶ eV cosmic rays (Gurevich et al., 2009; Gurevich and Zybin, 2005) or thermal runaway in leader and/or streamer tips (Celestin and Pasko, 2011; Colman et al., 2010; Moss et al., 2006). The runaway avalanche can result in a population of electrons with energies as high as tens of MeV inside the thundercloud (Shao et al., 2010; Dwyer et al., 2009; Gurevich et al., 1992). Whether relativistic electron beams can also occur above thunderclouds is currently not known (Östgaard et al., 2008; Inan, 2005; Smith et al., 2005; Roussel-Dupré and Gurevich, 1996; Fishman et al., 1994) even though transient narrow beams of
relativistic electrons are now routinely observed in space (Briggs et al., 2011; Cohen et al., 2010; Carlson et al., 2009; Dwyer et al., 2008). Relativistic electron beams are occasionally associated with sprites as inferred from radio remote sensing with low frequency radio signals from ~40–400 kHz (Füllekrug et al., 2010). These radio signals are analyzed here in detail to determine the height of relativistic electron beams above thunderclouds and to determine the associated charge transfer through the middle atmosphere.

2 The mesoscale convective system, lightning discharges and sprites

Thunderstorms and lightning activities in southern Europe have recently attracted special interest because numerous sprites are observed with video cameras above thunderclouds which regularly develop during the summer months from the beginning of May to mid-September (Soula et al., 2009; Neu- bert et al., 2008). For example, several thundercloud cells initiated over the Bay of Biscay and northern Spain in the late afternoon on 30 August 2008. The thunderstorms migrated northward through France and merged to an exceptional mesoscale convective system during the night until it decayed away in the late morning hours when reaching the English Channel which prevented an intrusion into Great Britain. The mesoscale convective system developed to a maximum area of ~400 × 400 km² as inferred from the cloud top infrared brightness temperature recorded on board the Meteosat Second Generation (MSG) satellite at 01:45 UTC (Fig. 1). The convective core of the thunderstorm reached a height of ~12 km with cloud top temperatures as low as ~−64 °C. Many negative lightning discharges cluster in the convective core which is surrounded by some positive lightning discharges reported by the French lightning detection system Météorage at the time when Meteosat scanned the electrified thundercloud from 01:50 UTC to 02:00 UTC.

One particular positive lightning discharge at 01:52:59.524 UTC (46.54°N, 0.48°E) with a peak current of ~47.1 kA reported by Météorage causes a sprite discharge above the thundercloud (Fig. 2) which is observed with a video camera located on the top of Pic du Midi in the French Pyrénées (42.9°N, 0.1°E) and Sant Vicenç de Castellet in North-Eastern Spain (41.7°N, 1.9°E). The sprite body is surrounded by three columniform sprites at distances up to ~20 km. The panels on the top and right hand side indicate the relative luminosity of the sprite which is inferred from the intensity of the image pixels.

The sprite luminosity curve is delayed by ~3–6 ms with respect to the causative positive lightning discharge (Fig. 3) as evidenced by high speed (20 kHz) recordings with a

Fig. 1. A meso-scale convective system is located over France and reaches an area of ~400 × 400 km² at 01:45 UTC on 31 August 2008 as inferred from the cloud top infrared brightness temperature. Many negative lightning discharges (circles) cluster in the ~12 km (Tmin = −64 °C) high convective core of the thunderstorm surrounded by some positive lightning discharges (crosses). One particular positive lightning discharge (red cross) at 01:52:59.524 UTC causes a sprite discharge above the thundercloud. The size of the red cross indicates the area covered by individual sprite elements.

Fig. 2. The main body of the sprite extends upwards from ~60–85 km height and the tendrils of the sprite extend downwards from ~60–50 km as triangulated by use of two video cameras from Pic du Midi in the French Pyrénées (42.9°N, 0.1°E) and Sant Vicenç de Castellet in North-Eastern Spain (41.7°N, 1.9°E). The sprite body is surrounded by three columniform sprites at distances up to ~20 km. The panels on the top and right hand side indicate the relative luminosity of the sprite which is inferred from the intensity of the image pixels.
3 Relativistic electron beams

Seven similar untypical broadband bursts of electromagnetic radiation following sprite-producing lightning discharges have been collected during sprite observations in the years 2008 and 2009 (Table 1). In these two years, a total of ∼140 sprite observations have been examined such that the occurrence rate of the electromagnetic radiation bursts is ∼5%. These bursts all occurred ∼2–9 ms after sprite-producing positive cloud-to-ground lightning discharges. The average spectrum of the seven bursts is relatively flat at frequencies <40 kHz when compared to the average spectrum of their parent lightning discharges (Fig. 4, left). This is consistent with simulations of virgin air high frequency breakdown.

The spectral amplitudes of the broadband electromagnetic radiation bursts are as small as ∼100 μV/m Hz−1/2 in the frequency range from ∼40–400 kHz, but still ∼1–2 orders of magnitude above an average background spectrum which is recorded 10 ms prior to the causative lightning discharges for reference (Fig. 4, left). This background spectrum exceeds the ultimate sensitivity limit of the wideband digital radio receiver ∼1–2 μV/m Hz−1/2 by ∼1 order of magnitude (Fig. 4, right). Superimposed on the sensitivity limit of the receiver are the omni-present low-frequency, or long-wave, radio transmitters in this frequency range, e.g., the United Kingdom radio clock at 60 kHz (the former Rugby MSF signal, now transmitted from Anthorn), France Inter at 162 kHz, Europe 1 at 183 kHz, and BBC radio 4 at 198 kHz. These radio transmitters have a typical narrow bandwidth of ∼10–12 kHz and they exhibit similar spectral amplitudes when compared to the observed broadband (∼40–400 kHz) radiation bursts from electron beams.

Numerical simulations of relativistic electron beams above thunderclouds predict theoretically broadband bursts of electromagnetic radiation with a flat spectrum following sprite-producing lightning discharges (Füllekrug et al., 2010) in agreement with our experimental observations. These model calculations solve kinetic particle equations on the microscopic scale and Maxwell’s equations on the macroscopic scale simultaneously. The causative lightning continuing current drains charge from the thundercloud which exposes the area above the thundercloud to a quasi-static electric field. When this electric field exceeds the runaway breakdown threshold and the continuous flux of high-energy secondary cosmic ray electrons (Daniel and Stephens, 1974, Fig. 2) supplies sufficient energetic seed electrons (Roussel-Dupré et al., 1998, p. 920, eq. in right col.), an electron avalanche above the thundercloud is initiated which quickly develops into an upward propagating relativistic electron beam. The beamed electrons partially discharge the lightning electric field and thereby gain a mean energy of ∼7 MeV with a spread of ∼6 MeV while carrying a total charge of ∼−10 mC upward as inferred from the comparison between the measured and simulated electron beam spectrum (Fig. 4, left and right). Impulsive changes of this downward...
Table 1. Observations of several electron beams associated with sprites. The date, time, and location of the causative lightning discharges are reported by the French lightning detection system Météorage. The distances between the lightning discharges and the wideband digital radio receiver in South-West England vary between $\sim 490$–$1060$ km. The electromagnetic radiation bursts from the lightning discharge and the consecutive electron beam exhibit time delays from $\sim 2$–$9$ ms. The time delays between the direct wave and the sky wave from the electron beam can only be determined for distances up to $\sim 500$ km as a result of the larger signal to noise ratio of the electromagnetic recordings. The observed time delays from $\sim 25$–$92$ $\mu$s indicate that all the electron beams are located at heights between $\sim 16$–$72$ km, well above the top of the thundercloud at $\sim 12$ km (Fig. 1). The last listed event at 01:29 UTC on 31 August was not associated with any sprite and it indicates that relativistic electron beams may also occur independently of sprites. The full range of techniques described in the methodology section was used during observations from 00:00–04:00 UTC on 31 August 2008.

<table>
<thead>
<tr>
<th>date</th>
<th>time/UTC</th>
<th>location</th>
<th>lightning peak current/kA</th>
<th>lightning/receiver distance/km</th>
<th>lightning/beam delay/ms</th>
<th>direct/sky wave delay/µs</th>
<th>height km</th>
</tr>
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<tr>
<td>31.08.2008</td>
<td>01:52:59.524482</td>
<td>0.4767° E, 46.5408° N</td>
<td>47.1</td>
<td>571</td>
<td>4.48/9</td>
<td>59/52</td>
<td>41/72</td>
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<tr>
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<td>02:18:26.265226</td>
<td>0.5489° E, 46.8654° N</td>
<td>73.1</td>
<td>540</td>
<td>8.0</td>
<td>92 16</td>
<td></td>
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<tr>
<td>31.08.2008</td>
<td>02:42:20.621219</td>
<td>0.3618° E, 47.3155° N</td>
<td>60.4</td>
<td>489</td>
<td>2.5</td>
<td>75 22</td>
<td></td>
</tr>
<tr>
<td>24.08.2009</td>
<td>21:04:24.032714</td>
<td>2.0992° E, 43.6166° N</td>
<td>42.2</td>
<td>920</td>
<td>2.0</td>
<td>– –</td>
<td></td>
</tr>
<tr>
<td>24.08.2009</td>
<td>21:18:11.336514</td>
<td>1.8931° E, 43.8959° N</td>
<td>37.9</td>
<td>886</td>
<td>7.0</td>
<td>– –</td>
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<tr>
<td>02.09.2009</td>
<td>02:09:56.601840</td>
<td>3.5730° E, 42.8089° N</td>
<td>52.0</td>
<td>1048</td>
<td>5.5</td>
<td>– –</td>
<td></td>
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<tr>
<td>02.09.2009</td>
<td>02:56:15.533414</td>
<td>4.3115° E, 42.9315° N</td>
<td>11.6</td>
<td>1060</td>
<td>2.0</td>
<td>– –</td>
<td></td>
</tr>
<tr>
<td>31.08.2008</td>
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<td>0.3979° E, 46.6625° N</td>
<td>63.1</td>
<td>557</td>
<td>5.5</td>
<td>60 43</td>
<td></td>
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</table>

(a) The average spectra of radiation from electron beams (red lines) and their causative lightning discharges (blue lines) on the left compare well with the theoretical simulations on the right. (a): The average spectrum of radiation produced by the seven measured electron beams (Table 1) is relatively flat at frequencies $<40$ kHz (red line) when compared with the average spectrum of their causative lightning discharges (blue line). The lightning and electron beam spectra are $\sim 1$–$2$ orders of magnitude larger than the average background spectrum (black line) which is recorded 10 ms prior to the causative lightning discharges for reference. (b): The radiation from the simulated electron beam is relatively flat at frequencies $<40$ kHz (red line) when compared with the simulated spectrum of the causative lightning discharge (blue line). The simulated spectra are $\sim 2$–$3$ orders of magnitude larger than the sensitivity limit of the wideband digital radio receiver $\sim 1$–$2$ $\mu$V m$^{-1}$ Hz$^{-1/2}$ which is inferred from time intervals without any bursts of electromagnetic radiation (black line). Superimposed on this ultimate reference spectrum are the omni-present low frequency, or long wave, radio transmitters, the strongest of which is here the BBC Radio 4 transmitter at 198 kHz, which exhibits spectral amplitudes comparable to the measured and simulated electron beam.

Fig. 4. The average spectra of radiation from electron beams (red lines) and their causative lightning discharges (blue lines) on the left compare well with the theoretical simulations on the right. (a): The average spectrum of radiation produced by the seven measured electron beams (Table 1) is relatively flat at frequencies $<40$ kHz (red line) when compared with the average spectrum of their causative lightning discharges (blue line). The lightning and electron beam spectra are $\sim 1$–$2$ orders of magnitude larger than the average background spectrum (black line) which is recorded 10 ms prior to the causative lightning discharges for reference. (b): The radiation from the simulated electron beam is relatively flat at frequencies $<40$ kHz (red line) when compared with the simulated spectrum of the causative lightning discharge (blue line). The simulated spectra are $\sim 2$–$3$ orders of magnitude larger than the sensitivity limit of the wideband digital radio receiver $\sim 1$–$2$ $\mu$V m$^{-1}$ Hz$^{-1/2}$ which is inferred from time intervals without any bursts of electromagnetic radiation (black line). Superimposed on this ultimate reference spectrum are the omni-present low frequency, or long wave, radio transmitters, the strongest of which is here the BBC Radio 4 transmitter at 198 kHz, which exhibits spectral amplitudes comparable to the measured and simulated electron beam.

Current from the electron beam $\sim 3 \times 10^{-3}$ Am$^{-2}$, i.e., its displacement current, results in the observed bursts of electromagnetic radiation a few ms after the causative lightning discharge.

The initial lightning discharge and the subsequent electron beam launch electromagnetic waves which can propagate on several paths from the source location to the radio receiver. These propagation paths can be inferred from a detailed analysis of the observed electromagnetic radiation bursts. The burst of electromagnetic radiation from the lightning discharge is composed of two consecutive pulses which are delayed by $\sim 102$ $\mu$s (Fig. 5, upper panel). The first pulse results
from the electromagnetic wave which propagates from the lightning discharge to the receiver along the surface of the Earth and is denoted ground wave. The second pulse results from the electromagnetic wave which is reflected from the ionosphere and is denoted sky wave. The time delay between the ground and sky wave determines the ionospheric reflection height (Jacobson et al., 2010) to be \( \sim 90 \) km for the known distance of \( \sim 571 \) km between the lightning discharge and the radio receiver as inferred from full wave radio wave propagation simulations with Finite-Difference-Time-Domain (FDTD) modeling. Electromagnetic waves with two or more ionospheric reflections are more strongly attenuated and they are not detected above the background level of the interfering radio transmitters (Smith et al., 2004).

The burst of electromagnetic radiation from the electron beam at \( \sim 4.4 \) ms is also composed of two consecutive pulses which are delayed by \( \sim 59 \) µs (Fig. 5, lower panel). The first pulse results from the direct wave which propagates at an angle to the receiver and is denoted direct wave. The second pulse results from the electromagnetic wave which is reflected from the ionosphere and is denoted sky wave. The time delay between the direct wave and the sky wave determines the height of the electron beam to be \( \sim 41 \) km for the known ionospheric height of \( \sim 90 \) km (Fig. 6). The electron beam is therefore located well above the thundercloud which terminates at an altitude of \( \sim 12 \) km. The burst of electromagnetic radiation at \( \sim 8.9 \) ms emanates from a height of \( \sim 72 \) km and could result from an independent second electron beam or from the same electron beam. In the latter case, the second radiation burst is produced by an acceleration of the electrons within the beam, i.e., a displacement current. The electron beam would then propagate with an average velocity of \( \sim 7000 \) km s\(^{-1}\) (\( \sim 0.023c \)) through the middle atmosphere as inferred from the difference between the two emission heights of \( \sim 31 \) km and the corresponding time delay \( \sim 4 \) ms. The heights of two more electron beams associated with other sprite discharges during the same night are determined to be \( \sim 16 \) km and \( \sim 22 \) km, both of which are well above the thundercloud.

4 Discussion

A sensitivity analysis of the height determination methodology reveals that a variation of the observed time delay by \( \sim 1 \) µs corresponds to a \( \sim 1 \) km vertical height variation such that the height of the electron beam can be determined in a very accurate way. If the observed time delay is interpreted as a variation in the distance between the source and the receiver, \( \sim 1 \) µs would correspond to a \( \sim 10 \) km distance difference, such that the source would be well outside the area of the mesoscale convective system. Another sensitivity analysis was performed to determine the optimum distance between the lightning discharge and the radio receiver which maximizes the signals of the ground and sky

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**Fig. 5.** The burst of electromagnetic radiation from the lightning discharge is composed of two consecutive pulses which are delayed by \( \sim 102 \) µs (upper panel). The first pulse results from the electromagnetic wave which propagates along the surface of the Earth to the receiver (ground wave). The second pulse results from the electromagnetic wave which is reflected from the ionosphere (sky wave). The burst of electromagnetic radiation from the electron beam is also composed of two consecutive pulses (lower panel). The short time delay of \( \sim 59 \) µs indicates that the electron beam is located above the thundercloud where it emits an electromagnetic wave which propagates at an angle to the receiver (direct wave).

**Fig. 6.** The time delay of \( \sim 102 \) µs between the ground wave and the sky wave from the lightning discharge indicates that the reflecting ionosphere is located at a height of \( \sim 90 \) km (upper panel). The time delay of \( \sim 59 \) µs between the direct wave and the sky wave from the electron beam indicates that the electron beam is located at a height of \( \sim 41 \) km (lower panel). The electron beam does not emit enough photons at visible wavelengths to exceed the sensitivity limit of the charge coupled device chip of the video camera. It is concluded that electron beams above thunderclouds can coincide with sprites but that they independently discharge lightning electric fields above thunderclouds.
wave. This distance was found to be \( \sim 300-400 \text{ km} \) such that at larger distances, it becomes increasingly difficult to distinguish the ground and sky wave from the background noise of the interfering radio transmitters. Finally, all lightning discharges with a peak current \( |I_p| > 30 \text{ kA} \) reported by Météorage from 00:00–04:00 UTC on 31 August 2008, have been investigated. Only one additional electromagnetic radiation burst with a flat spectrum from \( \sim 40-400 \text{ kHz} \) could be found which is similar to the three examples reported here. The burst follows \( \sim 5 \text{ ms} \) after two consecutive positive lightning discharges at 01:29:09.785 and 786 UTC (Table 1). This singular observation is indicative of a non-luminous relativistic electron beam without any sprite occurrence as evidenced by the two video cameras. The preliminary result suggests that lightning discharges may also cause relativistic electron beams above thunderclouds without producing sprites because the relativistic breakdown threshold is only \( \sim 1/10 \) of the conventional breakdown threshold. It is also possible that the observed electromagnetic radiation bursts are related to secondary sprite processes sometimes observed above thundercloud tops (Pasko, 2010; Marshall and Inan, 2007; Moudry, 2003). However, these processes remained sub-visible in the optical observations reported in this work, and it is not clear why such a hypothetical process would emit low frequency electromagnetic radiation while the sprite itself does not. It is interesting to note that the low frequency electromagnetic radiation reported here can be observed in space as a result of its transionospheric propagation (Füllekrug et al., 2011, 2009) and it may be associated with high frequency electromagnetic radiation (Parrot et al., 2008).

5 Summary

In summary, relativistic electron beams above thunderclouds have been detected in association with sprites. The electron beams can coincide with sprites but they are then located above the thundercloud and below the main body of sprites. The electron beams can also occur shortly after the sprite luminosity. Both results suggest that non-luminous electron beams and luminous sprites independently discharge lightning electric fields in the middle atmosphere. The observed relativistic electron beams above thunderclouds occur simultaneously with \( \sim 5 \% \) of all optically observed sprites and they are thus very rare. The small number of photons possibly emanating from non-luminous relativistic electron beams and sub-visible streamers may be identified in future studies by combining more sensitive optical observations with interferometric radio recordings to map the low-frequency radio sky above thunderclouds. However, the relativistic electron beams are a new form of impulsive energy transfer between thunderclouds and the middle atmosphere which need to be considered as a novel element in the global atmospheric electric circuit (Rycroft and Odzimek, 2010; Rycroft, 2006; Rycroft et al., 2000).

6 Methods

The Meteosat Second Generation satellite measures the thermal infrared radiation of the Earth from 10.5–12.5 \( \mu \text{m} \) with a Ritchey-Chrétien telescope connected to a HgCdTe detector. This passive imaging radiometer is calibrated with two black bodies at 290 K and 340 K and it scans the full Earth disk every 15 min. The optical camera system on Pic du Midi consists of a wide-angle, low-light, video camera and a high-speed photometer mounted on a pan-tilt unit. The camera is equipped with a 16 mm, \( f \) 1.40 lens with a field of view of \( 31^\circ \) and it has an exposure time of 40 ms for one image. The photometer counts photon emissions at a sampling rate of 20 kHz with a time resolution of 50 \( \mu \text{s} \). The camera system is synchronized to UTC with a Global Positioning Satellite (GPS) receiver with an absolute timing accuracy <1 ms. The optical camera system in Sant Vicенç de Castellet is a Watec 902H2 equipped with a 12 mm, \( f \) 0.8 lens with a field of view of \( 31^\circ \) and an exposure time of 20 ms. The French lightning detection system Météorage covers southwestern Europe and the western Mediterranean Sea. The radio detection system of 18 sensors reports cloud-to-ground lightning discharges with a detection efficiency of \( \sim 90 \% \), peak current accuracy of \( \sim 5 \% \), location accuracy \( <4 \text{ km} \), and timing accuracy \( <1 \text{ ms} \). The wideband digital radio receiver near Bath in the UK measures the vertical electric field strength by use of a capacitive probe with a sampling frequency of 1 MHz, frequency response from \( \sim 4 \text{ Hz} \) to \( \sim 400 \text{ kHz} \), amplitude resolution of \( \sim 35 \mu \text{V} \), and timing accuracy of \( \sim 12 \text{ ns} \). Numerical simulations of relativistic electron beams are performed with the transient, multimaterial, compressible, fluid dynamics code CAVEAT which was adapted to solve the electromagnetic equations and their effect on the electron and ion populations in a self-consistent way on the Los Alamos National Laboratory supercomputer facility. The Finite-Difference-Time-Domain (FDTD) method dynamically solves Maxwell’s equations on a spatial grid using equivalent vertical current sources representing the relativistic electron beam and the causative lightning discharge and propagates the full wave radio electromagnetic fields to the wideband digital radio receiver.

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